



Photon production from gluon fusion in heavy-ion collisions

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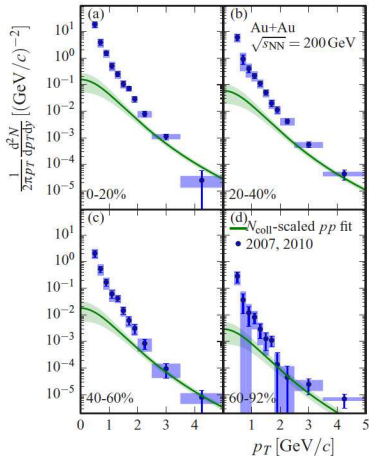
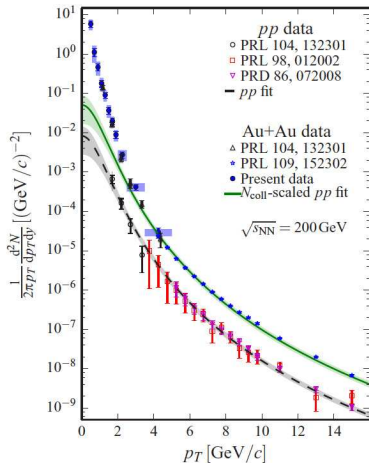
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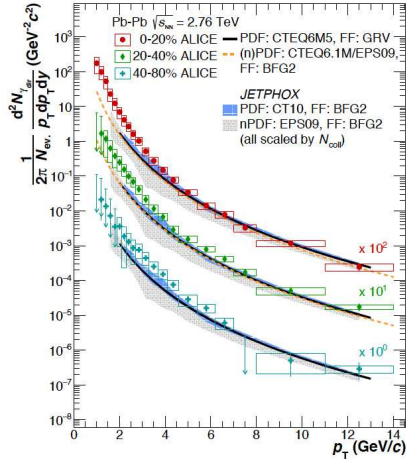
11th international workshop on High-pT Physics in the RHIC &
LHC Era, BNL April 2016



PHENIX, Phys. Rev. C 91, 064904 (2015)



ALICE, Phys. Lett. B **754**, 235-248 (2016)

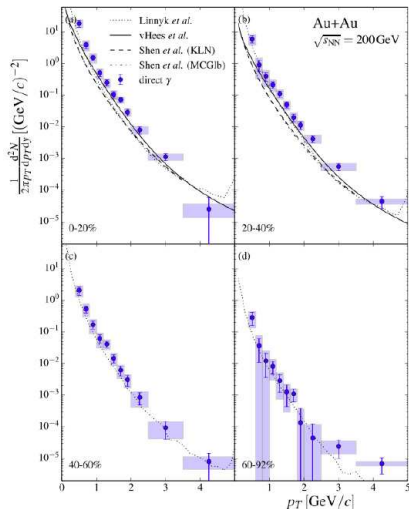


Outline

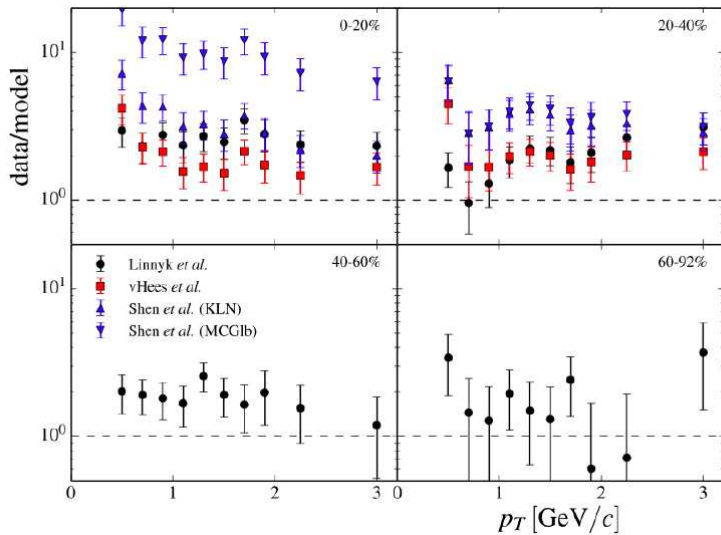
- ▶ Thermal photon excess as of 2014
- ▶ Current status: Data vs. models
- ▶ Magnetic field effects as a source of excess yield
- ▶ Gluon fusion into photons
- ▶ Conclusions

PHENIX compared to models as of 2014: Yield underestimated

- Transport model: Linnyk, Cassing, Bratkovskaya, Phys. Rev. C 89, 034908 (2014)
- Fireball model: van Hees, Gale, Rapp, Phys. Rev. C 84, 054906 (2011)
- Hydro model: Shen, Heinz, Paquet, Gale, Phys. Rev. C 89, 044910 (2014)

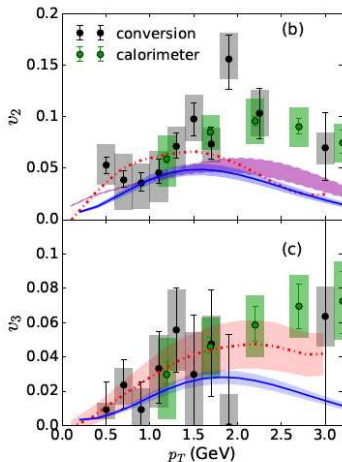
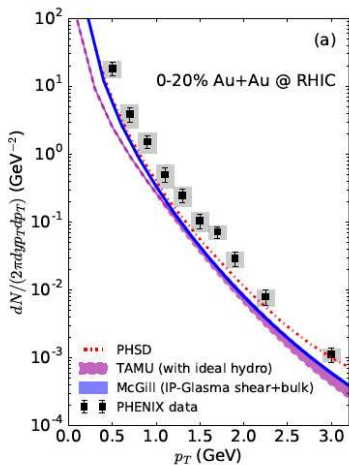


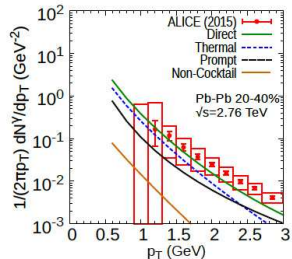
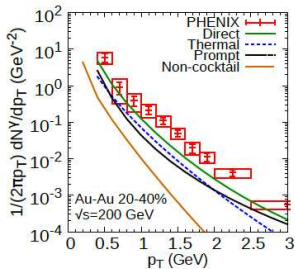
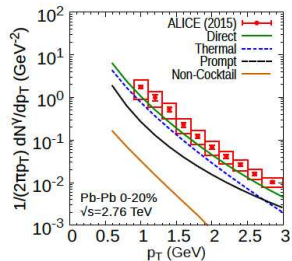
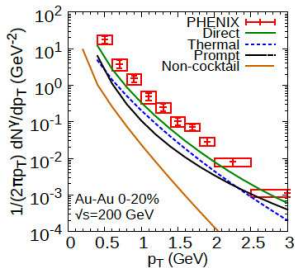
PHENIX compared to models as of 2014: Yield underestimated



PHENIX compared to models as of 2016: improvement

C. Shen e-Print: arXiv:1601.02563

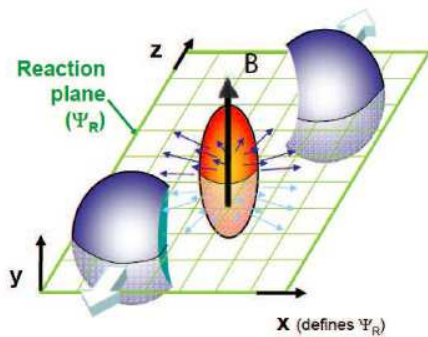


J.-F. Paquet *et al.*, arXiv:1509.06738

Summary of photon excess yield and flow:

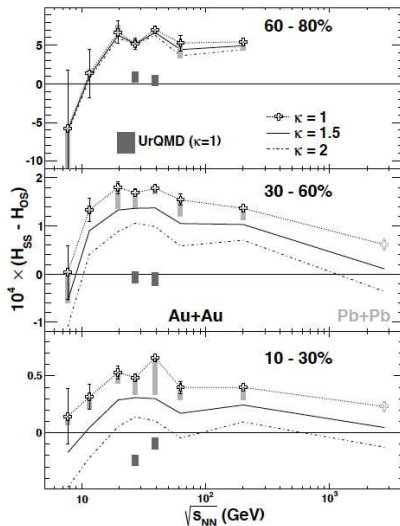
- ▶ Models have improved (including updated photon emission rates, viscosity, IP-Glasma)
- ▶ Models consistent with yield within the **(lower part of)** uncertainties
- ▶ v_2 still not well described
- ▶ v_3 better described than v_2

Magnetic field in heavy-ion collisions

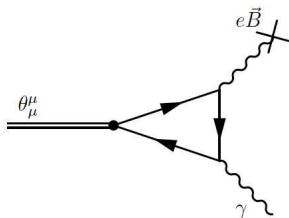


Charge separation along the magnetic field

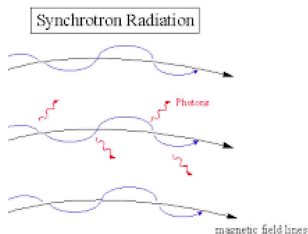
W. Broniowski, W. Florkowski, Phys. Rev. C 65, 024905 (2002)



Photons from magnetic field



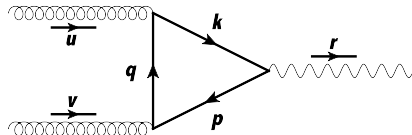
- ▶ Trace anomaly converts energy-momentum of gluon bulk into photons
G. Basar, D. Kharzeev, V. Skokov
Phys. Rev. Lett. **109**, 202303 (2012)



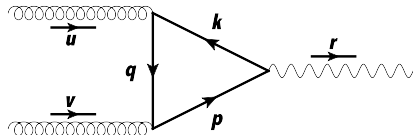
- ▶ Photon emission by quark synchrotron radiation
K. Tuchin, Phys. Rev. C **91**, 014902 (2015)

Photon production from gluon fusion

(a)



(b)



- ▶ Gluons are far more abundant than quarks at early times
- ▶ Largest magnetic field intensities and largest temperatures at early times of the collision
- ▶ Diagrams have same relative sign in the presence of magnetic field

Fermion propagator in magnetic field

The fermion propagator in coordinate space cannot longer be written as a simple Fourier transform of a momentum propagator

$$S(x, x') = \Phi(x, x') \int \frac{d^4 p}{(2\pi)^4} e^{-ip \cdot (x - x')} S(p)$$

where

$$\Phi(x, x') = \exp \left\{ iq_f \int_{x'}^x d\xi^\mu \left[A_\mu + \frac{1}{2} F_{\mu\nu} (\xi - x')^\nu \right] \right\}$$

is called the *phase factor*

Matrix element

$$\begin{aligned}
 \mathcal{M}^{(a)} &= - \int d^4x \int d^4y \int d^4z \int \frac{d^4p}{(2\pi)^4} \int \frac{d^4q}{(2\pi)^4} \int \frac{d^4k}{(2\pi)^4} \\
 &\times e^{-ip \cdot (y-x)} e^{-iq \cdot (z-y)} e^{-ik \cdot (x-z)} e^{-iu \cdot z} e^{-iv \cdot y} e^{ir \cdot x} \\
 &\times \text{Tr} \left[iq_f \gamma_\mu iS(k) ig \gamma_\alpha t^c iS(q) ig \gamma_\nu t^d iS(p) \right] \\
 &\times \Phi(x, y) \Phi(y, z) \Phi(z, x) \epsilon^{*\mu}(\lambda_r) \epsilon^\alpha(\lambda_u) \epsilon^\nu(\lambda_v).
 \end{aligned}$$

Fermion propagator in magnetic field: Intense field \implies Lowest Landau Level

The piece of the propagator $S(p)$ in momentum-space

$$\begin{aligned}
 iS(p) &= \int_0^\infty \frac{ds}{\cos(q_f Bs)} e^{is(p_\parallel^2 - p_\perp^2 \frac{\tan(q_f Bs)}{q_f Bs} - m_f^2 + i\epsilon)} \\
 &\times \left[(\cos(q_f Bs) + \gamma_1 \gamma_2 \sin(q_f Bs)) (m_f + \not{p}_\parallel) - \frac{\not{p}_\perp}{\cos(q_f Bs)} \right] \\
 &\xrightarrow{LLL} 2ie^{-\frac{p_\perp^2}{q_f B}} \frac{(\not{p}_\parallel + m_f)}{p_\parallel^2 - m_f^2} \left[\frac{1 - i\gamma_1 \gamma_2}{2} \right]
 \end{aligned}$$

The operator

$$\mathcal{O}_\parallel = [1 - i\gamma_1 \gamma_2] / 2$$

projects onto the longitudinal space. Therefore the **matrix element can be factorized into a product of transverse and longitudinal pieces**

Matrix element factorizes into product of transverse and longitudinal pieces

$$\begin{aligned}
 \mathcal{M}^{(a)} &= (2\pi)^4 \delta^4(r - v - u) \mathcal{M}_{\perp}^{(a)} \mathcal{M}_{\parallel}^{(a)} \\
 \mathcal{M}_{\perp}^{(a)} &= \left(\frac{4\pi}{q_f B} \right)^2 \int \frac{d^2 p_{\perp}}{(2\pi)^2} \int \frac{d^2 q_{\perp}}{(2\pi)^2} \int \frac{d^2 k_{\perp}}{(2\pi)^2} \\
 &\times e^{-\frac{k_{\perp}^2}{q_f B}} e^{-\frac{q_{\perp}^2}{q_f B}} e^{-\frac{p_{\perp}^2}{q_f B}} \prod_{i,j=1,2} e^{i \frac{2}{q_f B} \epsilon_{ij} (q-k+u)_i (q-p-v)_j} \\
 &= \left(\frac{q_f B}{12\pi} \right) e^{-\frac{(u+v)_{\perp}^2}{3q_f B}}, \\
 \mathcal{M}_{\parallel}^{(a)} &= -8 \left(\frac{q_f g^2 \delta^{cd}}{2} \right) \int \frac{d^2 p_{\parallel}}{(2\pi)^2} \int \frac{d^2 q_{\parallel}}{(2\pi)^2} \int \frac{d^2 k_{\parallel}}{(2\pi)^2} \\
 &\times (2\pi)^4 \delta^2 [(q-k+u)_{\parallel}] \delta^2 [(q-p-v)_{\parallel}] \epsilon^{*\mu}(\lambda_r) \\
 &\times \frac{\text{Tr} [\gamma_{\mu} \not{k}_{\parallel} \mathcal{O}_{\parallel} \gamma_{\alpha} \not{q}_{\parallel} \mathcal{O}_{\parallel} \gamma_{\nu} \not{p}_{\parallel} \mathcal{O}_{\parallel}]}{k_{\parallel}^2 q_{\parallel}^2 p_{\parallel}^2} \epsilon^{\alpha}(\lambda_u) \epsilon^{\nu}(\lambda_v)
 \end{aligned}$$

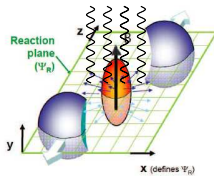
Early stages, gluons more abundant, quarks don't thermalize

- ▶ Hierarchy of energy scales $\sqrt{eB} > T > m$
- ▶ Early stages gluons far more abundant than quarks
- ▶ Assume quarks do not yet thermalize.
- ▶ $m_f = 0$ since in the absence of thermal corrections, the light-quark vacuum masses are negligible.
- ▶ The trace contains the product of up to twelve gamma matrices. Upon squaring and summing over polarizations, only a small piece survives

$$\begin{aligned} \text{Tr} [\gamma_\mu k_\parallel \mathcal{O}_\parallel \gamma_\alpha \not{q}_\parallel \mathcal{O}_\parallel \gamma_\nu \not{p}_\parallel \mathcal{O}_\parallel] &\longrightarrow k_{\parallel\nu} (p_{\parallel\mu} q_{\parallel\alpha} - p_{\parallel\alpha} q_{\parallel\mu}) \\ &+ k_{\parallel\mu} (p_{\parallel\nu} q_{\parallel\alpha} + p_{\parallel\alpha} q_{\parallel\nu}) \\ &+ k_{\parallel\alpha} (p_{\parallel\nu} q_{\parallel\mu} - p_{\parallel\mu} q_{\parallel\nu}) \end{aligned}$$

Kinematical simplifications

- ▶ For photons emitted at mid-rapidity, **momentum components along the reaction plane are small.**
- ▶ Reaction plane is perpendicular to the magnetic field then $r_{\perp} = (u + v)_{\perp} \simeq 0$.
- ▶ Focus on photons with small momentum $r_3 = (v + u)_3 \simeq 0$.
- ▶ Focus on describing emission of real photons $r^2 = (u + v)^2 = 0$.
- ▶ The main thermal effect on low momentum gluons is the development of a **thermal mass** $m_g \sim gT$.



Photon yield $m_f = 0$, thermal gluons, low p_t photons

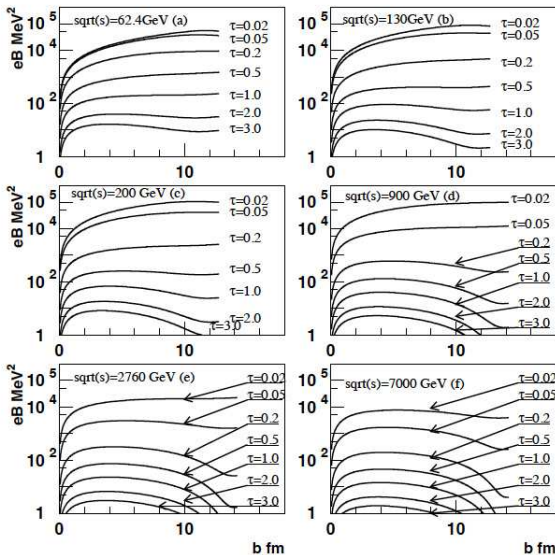
$$\frac{r_0 dN}{d^3r} = \frac{1}{2(2\pi)^3} \int \frac{d^3u}{2u_0(2\pi)^3} \int \frac{d^3v}{2v_0(2\pi)^3} \\ \times \sum_{\text{pol}, f} |\mathcal{M}|^2 n(u_0) n(v_0)$$

$$\frac{1}{2\pi N r_t} \frac{dN}{dr_t} = \frac{\left[\left(\frac{1}{3}\right)^4 e^{-2\frac{y_0^2 r_t^2}{eB}} + \left(\frac{2}{3}\right)^4 e^{-\frac{y_0^2 r_t^2}{eB}} \right] \frac{l(r_t/T)}{2\pi}}{\sqrt{3\pi eB/2} \left[\left(\frac{2}{3}\right)^{9/2} + \left(\frac{1}{3}\right)^{9/2} \right] \int_0^\infty dr_t l(r_t/T)}$$

$$l(z; \lambda) \equiv \int_0^z \frac{dx x^2 n\left(\sqrt{(z+x)^2 - (2x)^2}\right) n(x)}{\left(\sqrt{x^2 + \lambda^2}\right) \left(\sqrt{(z+x)^2 - (2x)^2 + \lambda^2}\right)}$$

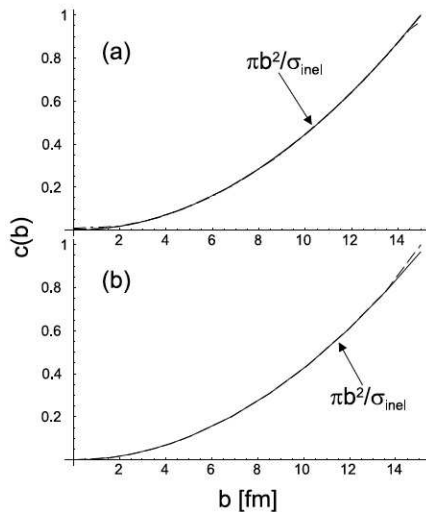
Magnetic field strength with time and impact parameter

L. Adamczyk, *et al.* (STAR COllaboration), Phys. Rev. Lett. **113**, 052302 (2014)



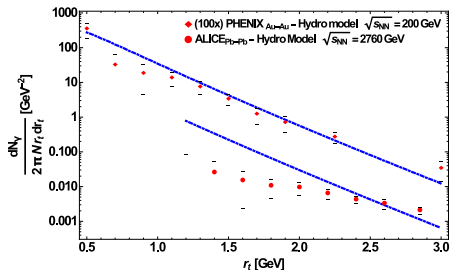
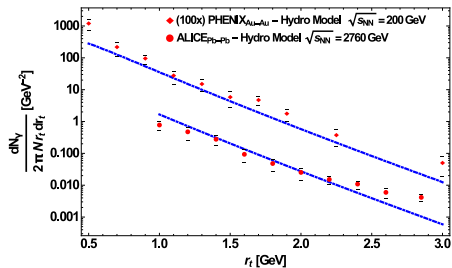
Centrality vs. impact parameter

W. Broniowski, W. Florkowski, Phys. Rev. C 65, 024905 (2002)



Excess photon yield from magnetic field induced gluon fusion compared to McGuil hydro

A.A., J. Castaño-Yepes, C. A. Dominguez, L. Hernández, arXiv:1604.02713



- ▶ $T = 300 \text{ MeV}$ for $\sqrt{s_{NN}} = 200 \text{ GeV}$ (RHIC)
- ▶ $T = 350 \text{ MeV}$ for $\sqrt{s_{NN}} = 2.76 \text{ TeV}$ (LHC)
- ▶ $g = 1$
- ▶ $0.5 \times 10^4 < eB/(\text{MeV})^2 < 10^5$ for RHIC
- ▶ $eB/(\text{MeV})^2 \simeq 10^4$ for LHC

Conclusions

- ▶ Magnetic fields originate observable phenomena in heavy-ion collisions
- ▶ Magnetic field induce processes otherwise now allowed such as emission of photons from gluon fusion
- ▶ Size of excess yield can be accounted for using reasonable values for eB , T and g .
- ▶ Magnetic fields naturally induce asymmetry (v_2).
- ▶ v_3 seems to be better described by models without magnetic field than v_2 .